

**<sup>24</sup>Na - Comments on evaluation of decay data  
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The first DDEP evaluation of <sup>24</sup>Na decay data was done by R.G. Helmer and E. Schönfeld in 2001 with a minor correction in 2004 (2004BeZR). The current evaluation has been completed in March 2014 with a literature cut-off by the same date.

### 1. DECAY SCHEME

The decay scheme of <sup>24</sup>Na is complete since the four levels populated in its β<sup>-</sup> decay are the only excited levels in <sup>24</sup>Mg below the decay energy. The spins, parities, and half-lives of the excited levels are from the evaluation by Firestone (2007Fi14).

### 2. NUCLEAR DATA

Q<sup>-</sup> value is from 2012 mass evaluation by Wang *et al.* (2012Wa38).

The recommended half-life of <sup>24</sup>Na is based on the experimental results given in Table 1.

**Table 1.** Experimental values of the <sup>24</sup>Na half-life (in hours)

| No. | Author(s) and year          | Reference | T <sub>1/2</sub> | Method and comments   |
|-----|-----------------------------|-----------|------------------|---|
| 1   | Wilson and Bishop (1949)    | 1949Wi10  | 14.90 (2)        | β counting  |
| 2   | Cobble and Atteberry (1950) | 1950Co69  | 15.10 (4)        | Ionization chamber  |
| 3   | Lockett and Thomas (1953)   | 1953Lo09  | 14.97 (2)        | Electroscope  |
| 4   | Tobailem (1955)             | 1955To07  | 14.90            | Ionization chamber; <i>omitted</i> - no uncertainty               |
| 5   | Campion and Merritt (1958)  | 1958Ca20  | 14.959 (10)      | Proportional counter  |
| 6   | Wolf(1960)                  | 1960Wo07  | 14.953 (13)      | Ionization chamber  |
| 7   | Wyatt et al. (1961)         | 1961Wy01  | 15.05 (2)        | Ionization chamber; <i>omitted</i> as superseded by 11            |
| 8   | Monahan et al. (1962)       | 1962Mo21  | 15.04 (5)        | Ionization chamber  |
| 9   | Lagoutine et al. (1968)     | 1968La10  | 15.00 (2)        | Ionization chamber; <i>omitted</i> as superseded by 19            |
| 10  | Kemeny (1969)               | 1969Ke14  | 15.16 (5)        | Ionization chamber; <i>omitted</i> - no background subtraction    |
| 11  | Emery et al. (1972)         | 1972Em01  | 15.030 (3)       | Ionization chamber; <i>omitted</i> as outlier                     |
| 12  | Chakraborty (1974)          | 1974Ch25  | 14.969 (12)      | Ionization chamber; average of 6 values with external uncertainty |

|                          |                                  |          |                   |   |
|--------------------------|----------------------------------|----------|-------------------|---|
| 13                       | Genz et al. (1976)               | 1976Ge06 | 15.09 (6)         | Ionization chamber  |
| 14                       | Davis et al. (1978)              | 1978Da21 | 15.010 (28)       | Ionization chamber  |
| 15                       | Houtermans et al. (1980)         | 1980Ho17 | 14.9590 (12)      | 4 $\pi\gamma$ ionization chamber  |
| 16                       | Muckenheim et al. (1980)         | 1980Mu11 | 14.964 (15)       | Ge(Li) detector   |
| 17                       | Rutledge et al. (1980)           | 1980RuZY | 14.965 (10)       | NaI(Tl) detector;<br><i>omitted</i> as superseded by 18                 |
| 18                       | Rutledge et al. (1982)           | 1982RuZV | 14.965 (10)       | NaI(Tl) detector  |
| 19                       | Lagoutine, Legrand (1982)        | 1982La25 | 14.956 (3)        | Ionization chamber  |
| 20                       | Hoppes et al. (1982)             | 1982HoZJ | 14.951 (3)        | 4 $\pi\gamma$ ionization chamber;<br><i>omitted</i> as superseded by 24 |
| 21                       | Walz et al. (1983)               | 1983Wa26 | 14.9575 (28)      | 4 $\pi\gamma$ ionization chamber  |
| 22                       | Abzouzi et al. (1989)            | 1989Ab05 | 15.027 (2)        | $\gamma$ counting, NaI(Tl), Ge-detectors;<br><i>omitted</i> as outlier  |
| 23                       | Bode et al. (1991)               | 1991Bo34 | 14.90 (2)         | $\gamma$ counting, a well-type NaI(Tl), Ge- detectors                   |
| 24                       | Unterweger <i>et al.</i> (1992)  | 1992Un01 | 14.9512 (32)      | 4 $\pi\gamma$ ionization chamber;<br><i>omitted</i> as superseded by 26 |
| 25                       | Mignonsin (1994)                 | 1994Mi03 | 14.86 (12)        | $\gamma$ counting, GeLi detector  |
| 26                       | Unterweger <i>et al.</i> (2002)  | 2002Un02 | 14.951(3)         | 4 $\pi\gamma$ ionization chamber;<br><i>omitted</i> as superseded by 27 |
| 27                       | Unterweger and Lindstrom (2004)  | 2004Un01 | 14.9512 (32)      | 4 $\pi\gamma$ ionization chamber;<br><i>omitted</i> as superseded by 29 |
| 28                       | Lindstrom <i>et al.</i> (2005)   | 2005Li66 | 14.959 (9)        | $\gamma$ counting, Ge detector, 4 $\pi\gamma$ ionization chamber        |
| 29                       | Fitzgerald (2012)                | 2012Fi12 | 14.951 (3)        | 4 $\pi\gamma$ ionization chamber;<br><i>omitted</i> as superseded by 30 |
| 30                       | Unterweger and Fitzgerald (2014) | 2014Un01 | 14.955 (7)        | 4 $\pi\gamma$ ionization chamber  |
| <b>Recommended value</b> |                                  |          | <b>14.958 (2)</b> | <b>LWM</b>  |

Value 4 has been omitted as it is reported without uncertainty. Value 10 has been omitted as it was obtained without background subtraction. The values 7, 9, 17, 20, 24, 26, 27, 29 were not used because they were replaced ultimately by later results of the same laboratory.

In the final weighted average, the values of 11 (1972Em01) and 22 (1989Ab05) have been omitted because they are outliers; both are over 30  $\sigma$  from the adopted value. If these values are included, the reduced  $\chi^2 / (\chi^2)_{\text{crit}}$  value is about 70/2. For the 18 values included, the LWEIGHT computer program using the limitation of relative statistical weight method (LWM)

has increased the uncertainty of the value 15 (1980Ho17) from 0.0012 to 0.0018 in order to reduce its relative weight from 69% to 50%. In addition to this relative weight, those of the values of 1982La25 and 1983Wa34 are 17% and 20%.

For the final weighted average of 14.958 the internal uncertainty is 0.0013, the reduced- $\chi^2 / (\chi^2)_{\text{crit}}$  value is 2.6/2.0, and the external uncertainty is 0.0020. The LWEIGHT program has chosen the weighted average with the external uncertainty.

The recommended value of <sup>24</sup>Na half-life is **14.958 (2) hours**. It can be compared with earlier evaluations of 14.9574 (20) h (2004BeZR) and 14.997 (12) h (2007Fi14).

## 2.1. Beta Transitions

The energies of  $\beta^-$  transitions have been obtained using the  $Q^-$  value and the <sup>24</sup>Mg level energies adopted from 2007Fi14 (Table 2). The probabilities of  $\beta^-$ -transitions  $P_{\beta^-}$  have been deduced from the  $P(\gamma+ce)$  balance at each level of <sup>24</sup>Mg.

**Table 2.** <sup>24</sup>Mg levels populated in <sup>24</sup>Na  $\beta^-$ -decay

| Level | Energy (keV)  | Multipolarity | Half-life   | $\beta^-$ -transition probability (%) |
|-------|---------------|---------------|-------------|---------------------------------------|
| 0     | 0             | 0+            | Stable      | $< 5 \times 10^{-10}$                 |
| 1     | 1368.672 (5)  | 2+            | 1.33 (6) ps | 0.003 (2)                             |
| 2     | 4122.889 (12) | 4+            | 22 (2) fs   | 99.930 (3)                            |
| 3     | 4238.24 (3)   | 2+            | 41 (4) fs   | < 0.002                               |
| 4     | 5235.12 (4)   | 3+            | 61 (7) fs   | 0.066 (3)                             |

Below the nine experimentally determined values of the  $\beta^-$ -end-point energy (in keV) for the transition to the 4123-keV level are compared with the value derived from the Q value.

| Energy, keV | Reference | Author(s) and year           |
|-------------|-----------|------------------------------|
| 1394 (4)    | 1957Po36  | Porter et al. (1957)         |
| 1389 (4)    | 1958Da10  | Daniel (1958)                |
| 1389 (2)    | 1961De25  | Depommier and Chabre (1961)  |
| 1395        | 1963Pa20  | Paul et al. (1963)           |
| 1393 (3)    | 1964Le09  | Lehmann (1964)               |
| 1394 (2)    | 1965Be24  | Beekhuis and De Waard (1965) |
| 1389.2 (5)  | 1969Bo48  | Booij et al. (1969)          |

|                    |   |                    |
|--------------------|---|--------------------|
| 1389 (2)           | 1972Gi17                                | Gils et al. (1972) |
| 1390 (1)           | 1976Ge06                                | Genz et al. (1976) |
| <b>1392.72 (4)</b> | <b>Q - E<sub>level</sub> (4123 keV)</b> |                    |

The measured and adopted (calculated) probabilities (in %) of the  $\beta^-$  - transitions are:

| Level Energy<br>(keV) | 1950Gr01<br>Grant (1950) | 1951Tu12<br>Turner (1951) | <b>Adopted<br/><math>\beta^-</math> - transition probabilities</b> |
|-----------------------|--------------------------|---------------------------|--|
| 5235                  |                          |                           | <b>0.066 (3)</b>   |
| 4238                  |                          |                           | <b>0.001 (1)</b>   |
| 4123                  | 100                      | 100                       | <b>99.930 (3)</b>  |
| 1369                  | < 0.01                   | 0.003                     | <b>0.003 (2)</b>   |
| 0                     |                          |                           | <b>0</b>   |

The 4<sup>th</sup> forbidden  $\beta^-$ -branch to the ground state has not been observed. As discussed by Helmer and Schönfeld in 2004BeZR, from the experimental limit on the number of counts in the  $\beta^-$ -spectrum above 4140 keV, the authors of 1951Tu12 give  $\lg ft > 15.1$ . The  $\lg ft$  systematics of 1998Si17 lists four decays of this type with  $\lg ft$  values of 22.5 to 24.3. Since this is a very small set of values, for <sup>24</sup>Na the lower limit of the  $\lg ft$  can be taken to be 20, which corresponds to the transition probability to the ground state  $P_{\beta^-}(0) < 5 \times 10^{-10}$  %; hence no direct transition to the ground state has been adopted.

The  $\beta^-$ -branch to the 4238 keV level is a 2<sup>nd</sup> forbidden transition and the  $\lg ft$  systematics (1998Si17) give  $\lg ft > 10.6$  which corresponds to the transition probability to the 4238 keV level  $P_{\beta^-}(4238) < 0.002$  %; a value of 0.001 (1) % has been adopted. The negligible probability of this transition is supported by the adopted decay scheme for which the intensity of the 997-keV  $\gamma$ -ray feeding this level is more [0.00145 (25)] than that depopulating it [0.0011 (1) = 0.00025 (3) + 0.00084 (10)]. An unobserved  $\gamma$ -ray of 116 keV also theoretically depopulating this level could improve the intensity balance.

Therefore, no direct  $\beta^-$ -transition to the ground state and a transition probability of 0.001 (1) % to the level of 4238 keV, are present in the adopted <sup>24</sup>Na decay scheme.

No direct measurements are reported for the  $\beta^-$ -transitions to the 4238- and 5235-keV levels. The adopted value for the transition to the 1369-keV level is compatible with the measurement of 1951Tu12 who gave no uncertainty. The adopted value for the transition to the

5235-keV level was calculated from probabilities of the two de-exciting  $\gamma$ -rays and their electron and pair conversion.

## 2.2. Gamma Transitions and Internal Conversion Coefficients

The adopted ICC(s) and  $\alpha_{\pi}$  are the theoretical values interpolated by the BrIcc computer program (2008Ki07) from the tables of Band et al. (2002Ba85), accepting the “frozen orbital (no hole)” approximation. The multipolarities and mixing ratios  $\delta$  have been taken from 2007Fi14.

The transition probabilities of the 3866- and 4238-keV  $\gamma$ -rays have been determined from the following measurements of the  $\gamma$ -ray emission probabilities:

| Reference            | Author(s) and year        | P $\gamma$ (3866 keV) (%) | P $\gamma$ (4238 keV) (%) |
|----------------------|---------------------------|---------------------------|---------------------------|
| 1960Ar10             | Artamonova et al. (1960)  | 0.09 (2)                  | 0.0015 (5)                |
| 1962Mo21             | Monahan et al. (1962)     | 0.075 (20)                | 0.008 (3)                 |
| 1968Va06             | van Klinken et al. (1968) | 0.063 (6)                 |                           |
| 1970Le12             | Lebowitz (1970)           | 0.0489 (25)               | < 0.0033                  |
| 1972Ra21             | Raman et al. (1972)       | 0.061 (5)                 | 0.00084 (10)              |
| 2003Ep02             | Epp and Griffin (2003)    | 0.067 (2)                 | 0.00085 (39)              |
| <b>Adopted value</b> |                           | <b>0.066 (2)</b>          | <b>0.00084 (10)</b>       |

For the 3866-keV  $\gamma$ -ray, the value of 1970Le12 has been omitted because this value is too low (outlier). If this value is included, the reduced  $\chi^2 / (\chi^2)_{\text{crit}}$  value is about 7/3. For the 5 values included, a weighted average is 0.0662 with an internal uncertainty of 0.0018 and external uncertainty of 0.0016, the reduced-  $\chi^2 / (\chi^2)_{\text{crit}}$  value is 0.8/3.3. The LWEIGHT code has chosen the weighted average and internal uncertainty. The smallest experimental uncertainty is 0.002. Thus, the adopted value of P $\gamma$  (3866 keV) is 0.066 (2) %.

For the 4238-keV  $\gamma$ -ray, the adopted value is the average of three values (1962Mo21, 1972Ra21, 2003Ep02) which is consistent with the limit of 1970Le12.

The 997- and 2869-keV  $\gamma$ -ray transitions are not observed in <sup>24</sup>Na decay, but their emission probabilities can be deduced from the relative probabilities in other decays or reactions. The transition probability of 997-keV  $\gamma$ -ray was calculated from the measured P $\gamma$ (997)/P $\gamma$ (3866) ratio. For this ratio, the measured values are:

| P $\gamma$ (997)/P $\gamma$ (3866) | Reference | Author(s) and year                                      |
|------------------------------------|-----------|---|
| 0.017 (5)                          | 1972Me09  | Meyer et al. (1972) from <sup>23</sup> Na(p, $\gamma$ ) |

|                  |                      |   |
|------------------|----------------------|---|
| 0.019 (2)        | 1973Le15             | Leccia et al. (1973) from <sup>23</sup> Na(p,γ)         |
| 0.015 (3)        | 1975Bo43             | Boydell et al. (1975) from <sup>23</sup> Na(p,γ)        |
| 0.0260 (17)      | 1981Wa07             | Warburton et al. (1981) from <sup>24</sup> Al ε - decay |
| 0.030 (4)        | 1990En02             | Endt et al. (1990) from <sup>23</sup> Na(p,γ)           |
| <b>0.022 (4)</b> | <b>Adopted value</b> |   |

The adopted value is the weighted average value of 0.022, a reduced-  $\chi^2 / (\chi^2)_{\text{crit}}$  value is 4.6/3.3. The LRSW method increases the final uncertainty to 0.004 to include the most precise value of 0.0260. With the above value of P $\gamma$ (3866), we obtain P $\gamma$ (997) = 0.00145 (25).

The ratio P $\gamma$ (2869)/P $\gamma$ (4238) has been measured as follows:

| P $\gamma$ (2869)/P $\gamma$ (4238) | Reference            | Author(s) and year                                    |
|-------------------------------------|----------------------|---|
| 0.30 (3)                            | 1972Me09             | Meyer et al. (1972) from <sup>23</sup> Na(p,γ)        |
| 0.30 (3)                            | 1972Ra21             | Raman et al. (1972) from <sup>23</sup> Mg(n,n'γ)      |
| 0.299 (15)                          | 1973Le15             | Leccia et al. (1973) from <sup>23</sup> Na(p,γ)       |
| 0.267 (7)                           | 1973Br16             | Branford (1973) from <sup>23</sup> Na(p,γ)            |
| 0.299 (19)                          | 1975Bo43             | Boydell (1975) from <sup>23</sup> Na(p,γ)             |
| 0.304 (19)                          | 1981Wa07             | Warburton et al. (1981) from <sup>24</sup> Al ε decay |
| <b>0.300 (15)</b>                   | <b>Adopted value</b> |   |

The value of 1973Br16 has been omitted as an outlier. For the final weighted average of 0.3003 the internal uncertainty is 0.0090, the reduced-  $\chi^2 / (\chi^2)_{\text{crit}}$  value is 0.01/3.3, and the external uncertainty is 0.0010. The LWEIGHT program has chosen the weighted average with the internal uncertainty. The smallest experimental uncertainty is 0.015. The adopted value is 0.300 (15).

With the above adopted value of 0.00084 (10) for P $\gamma$ (4238), one obtains P $\gamma$ (2869) = 0.00025 (3).

If there are no direct feeding the ground state by  $\beta^-$ -decay or the unobserved  $\gamma$  transitions of 4123 and 5235 keV, T $\gamma$ (1369) = 100 – T $\gamma$ (4237) = 99.99916 (10) where T $\gamma$  = P $\gamma$  (1.0 +  $\alpha_T$  +  $\alpha_\pi$ ). Upper limits for transition intensities of the 4123- and 5235-keV  $\gamma$ -rays can be determined from the ratios measured by 1981Wa07: P $\gamma$ (4123)/P $\gamma$ (2754) < 0.00001, or P $\gamma$ (4122) < 0.001 and P $\gamma$ (5235)/P $\gamma$ (3866) < 0.004, so P $\gamma$ (5235) < 0.00023 and by 1972Ra21 and 1967En05 which give P $\gamma$ (4123) < 0.0009 and P $\gamma$ (5235) < 0.00002. If the 4123- and 5235-keV transitions have

intensities equal to the latter upper limits, the value of  $T\gamma(1369)$  would reduce from 99.99916 to 99.9983. Since it is unlikely that these two values will be at the limits, we have adopted the value of  **$T\gamma(1369) = 99.9990$  (3)** (see also 2004BeZR) and from here  **$P\gamma(1369) = 99.9934$  (5) %**.

The 1112-keV transition between the 5235- and 4123-keV levels has not been observed in <sup>24</sup>Na decay. In the <sup>24</sup>Al decay, the authors of 1981Wa07 have found an upper limit of the ratio  $P\gamma(1112)/P\gamma(3866) < 0.007$  which yields the value of  $P\gamma(1112) < 0.0004$ .

The transition probability of the 2754-keV  $\gamma$ -ray has been calculated from the balance condition  $T\gamma(2754) = T\gamma(1369) - [T\gamma(2869) + T\gamma(3866) + P\beta^-(4147)]$ . This yields  **$T\gamma(2754) = 99.9990$  (3) - 0.069 (3) = 99.930 (3) %**, which gives  **$P\gamma(2754) = 99.862$  (3) %**.

From the intensity balance at the 4238-keV level, for a possible depopulating  $\gamma$ -ray of 116 keV,  $P\gamma(116) = 0.0004$  (3) +  $I\gamma(4238)$ . Since this  $\gamma$ -ray has not been observed, it is omitted from the scheme.

The internal-pair formation coefficients ( $\alpha_\pi$ ) for the 1369- and 2754-keV  $\gamma$ -rays calculated by the BrIcc code (2008Ki07) can be compared with measured values:

| $\alpha_\pi$ (1369 keV) | $\alpha_\pi$ (2754 keV) | Reference                   |
|-------------------------|-------------------------|-----------------------------|
| 0.00116 (10)            |                         | 1949Ra01                    |
| 0.00076 (19)            |                         | 1950Mi82                    |
| 0.00067 (10)            |                         | 1951Cl50                    |
| 0.00006 (1)             | 0.00071 (2)             | 1952Bl53                    |
| 0.00003                 | 0.00080                 | 1952Sl52                    |
| <b>0.0000463 (7)</b>    | <b>0.000675 (10)</b>    | <b>adopted (calculated)</b> |

### 3. ATOMIC DATA

The atomic data (fluorescence yields, X-ray energies and relative probabilities, and Auger electrons energies and relative probabilities) have been deduced by using the SAISINUC software.

### 4. ELECTRON EMISSIONS

The energies of the conversion electrons and the energies of the electrons from the internal-pair creation process have been obtained from the gamma-ray transition energies. The absolute emission probabilities of the conversion electrons have been deduced using recommended  $P\gamma$ , ICC and  $\alpha_\pi$  values.

The absolute emission probabilities of K and L Auger electrons have been calculated using the EMISSION computer program.

$\beta^-$  average energies have been calculated using the LOGFT computer program.

## 5. PHOTON EMISSIONS

### 5.1. X-ray Emissions

The absolute emission probabilities of Mg KX- and LX- rays have been calculated using the EMISSION computer program.

### 5.2. Gamma ray emissions

The energies of gamma rays in <sup>24</sup>Mg have been obtained from the adopted level energies using the SAISINUC code.

The number of photons per disintegration was deduced as described in sect. 2.2.

## 6. ENERGY CONSERVATION

The total average energy of 5515.63 (12) keV, for one disintegration, calculated from the current evaluated data corresponds very well to the available energy of 5515.61 (4) keV ( $Q^-$ ) from the mass tables (2012Wa38) confirming the consistency of the decay scheme and the reliability of this evaluation.

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